

STATE RESEARCH CENTER OF RUSSIA INSTITUTE FOR HIGH ENERGY PHYSICS

IHEP 97-49

A.A. Derevschikov, Yu.A. Matulenko, V.L. Rykov¹, K.E. Shestermanov, A.N. Vasiliev, L.V. Alekseeva², L.V. Nogach²

SIMULATION STUDY OF γ/π^{o} -SEPARATION AND e/h-REJECTION POWER IN THE STAR BARREL ELECTROMAGNETIC CALORIMETER WITH THE GASEOUS SHOWER MAXIMUM DETECTOR

Protvino 1997

¹Current address: Wayne State University, Detroit, MI 48201, USA.

²Graduate student from Moscow State University.

Abstract

Derevschikov A.A., Matulenko Yu.A., Rykov V.L. et al. Simulation Study of γ/π^{o} -Separation and e/h-Rejection Power in the STAR Barrel Electromagnetic Calorimeter with the Gaseous Shower Maximum Detector: IHEP Preprint 97-49. – Protvino, 1997. – p. 14, figs. 9, refs.: 13.

The γ/π^{o} -separation has been studied using a gaseous Shower Maximum Detector (SMD) positioned at four different locations inside the STAR Electromagnetic Calorimeter (EMC), $3X_{o}$, $4X_{o}$, $5X_{o}$ and $7X_{o}$, for various energies of the incident particles in the range from 3 to 25 GeV.

Also the electron-hadron rejection power of the STAR EMC with the gaseous SMD has been evaluated for two SMD positions within the EMC stack at the energies from 2 to 30 GeV.

Аннотация

Деревщиков А.А., Матуленко Ю.А., Рыков В.Л. и др. Расчеты эффективности γ/π^{o} разделения и e/h-реакции в центральном электромагнитном калориметре и детекторе максимума ливня установки STAR: Препринт ИФВЭ 97-49. – Протвино, 1997. – 14 с., 9 рис., библиогр.: 13.

Методом Монте-Карло изучалась возможность разделения π^{o} -мезонов и одиночных γ квантов с помощью газового Детектора Максимума Ливня (ДМЛ) при его различных положениях внутри Электромагнитного Калориметра (ЭК) ($3X_o$, $4X_o$, $5X_o$ и $7X_o$) в диапазоне энергий падающих частиц от 3 до 25 ГэВ.

Также были проведены расчеты степени режекции электрон/адрон для энергий от 2 до 30 ГэВ и для двух положений ДМЛ (3X_o и 5X_o) внутри ЭК.

© State Research Center of Russia Institute for High Energy Physics, 1997

Introduction

In this paper we provide the estimates of how well the STAR Barrel Electromagnetic Calorimeter (EMC) with the gaseous Shower Maximum Detector (SMD) can distinguish single photons from the double-photon hits originated from π^0 -decay which are the main background for the study of direct photon production and asymmetries. The similar study for the scintillator strip/fiber SMD has been accomplished earlier and published in Ref. [1].

We have also estimated the EMC/SMD rejection power for e/π -separation in the energy range from 1 to 30 GeV which is important for both the Heavy-Ion and Spin Physics programs in STAR at RHIC.

The STAR is one of two large RHIC's detectors which, along with its rich capabilities for Heavy Ion Physics, will also carry out practically the entire Spin Program with polarized protons colliding at RHIC. Particularly, as it has been suggested in the RHIC Spin Proposal [2], the information on spin-dependent gluon structure functions will be obtained by measuring the asymmetries of direct photons and jet production at $P_T \geq 10 \text{ GeV/c}$ as well as by studying the production and decay asymmetries (including parity violation) for $W^{\pm} \to e^{\pm}\nu$ and $Z^0 \to e^+e^-$.

Single photons and dileptons are also regarded as the "penetrating" probes for Quark-Gluon Plasma (QGP) because they are produced at the early stage of the collision and are not affected by the subsequent hadronization of the system [3]. However, the region of interest here is at much lower $P_T \simeq 1-7$ GeV/c compared to polarized proton collisions. This, along with the much higher multiplicity of heavy-ion events, makes the task even more challenging.

At the stage of simulations presented here, we don't take into account the complications arising from the high multiplicity. We fully understand, however that, because of this simplification, our estimates for γ/π^o -separation at low P_T can be considered as rather a very "remote scouting" of the problem than its direct attack.

As it has been suggested in Ref. [4], the recognition of direct photons at $P_T \geq 7$ -10 GeV/c is also interesting for the exploration of Jet Quenching mechanism in heavy ion collisions. The quantitative information about this phenomenon can be obtained by measuring the high- P_T hadron spectra in the events tagged with the high- P_T direct photons. The problem of high multiplicity of heavy ion collisions seems, to some extent, less significant here than at lower P_T 's. However, these data are not expected to be among the first results from Au-Au collisions at RHIC since quite a high integrated luminosity is required for the accumulation of a statistically sufficient sample of events.

The results presented below have been obtained, using GEANT-3.21 code. The GEANT parameters for simulating the electromagnetic showers in the EMC/SMD were chosen, following the recommendation of Ref. [5]: ISTRA = 1, ILOSS = 1; with the thresholds of

- 500 KeV for electrons/positrons in Pb and Scintillator material;

-80 KeV for γ -quanta in Pb and Sc;

- 10 KeV for e and γ in the gaseous SMD.

To simulate the hadron showers, GHEISHA code have been used.

Usually, a sample of 500 events was generated at each SMD position in the EMC stack for each energy.

It should be noticed that, in these simulations, the energy depositions in active media (scintillator layers and SMD gas mixture) and only their fluctuations only taken were into account. All other effects (statistics of photoelectrons in the EMC, charge collection in the SMD, EMC and SMD segmentations, edge effects, etc.) were ignored which is apparently one more major simplification of this stage simulations. We suppose to include these effects at the next stage along with the GEANT parameters better "tuned up" to the EMC/SMD test beam data which are available now.

1. The Geometry of the Electromagnetic Calorimeter and Shower Maximum Detector

The EMC geometry used in these simulations was the non-segmented block of Scintillator-Lead calorimeter with the cross-section 70×70 cm² and the gaseous SMD placed inside the stack. The calorimeter stack consisted of 20 sandwich-type layers of the thickness approximately equal to 0.9 of radiation length (X_o) each. In turn, each layer consisted of 4 mm thick scintillator tile plus 1.3 mm G10 plate plus 5 mm thick Pb plate. The 19 mm thick Aluminum plate were placed in front of the stack, exactly as it's supposed to be in the real STAR EMC.

In our simulations, the "CDF-type" [6] SMD placed after the 3d — 7th EMC layer, was actually the multi-wire proportional chamber enclosed into the Aluminum flat box with the 2 mm thick front plate and 4 mm thick back plate. Just behind the front plate, the 4 mm thick G10 printed circuit board with the cathode strips at its surface was located. The 6 mm thick ionization sensitive volume between this board and Al back plate was filled with the 70% Ar plus 30% CO_2 gas mixture.

The readout from SMD in Y-direction was from the anode wires 7 mm apart. The readout in the orthogonal, X-direction, was from 14 mm wide signal cathode strips¹.

The initial particles from the source (γ , electrons, and $\pi^{0,-}$) were directed toward the EMC block central area of the size $10 \times 10 \, cm^2$, perpendicularly to its front plate. The source of incident particles was located at the 220 cm distance from the EMC front plate.

2. SMD Characteristics

In 1995, the IHEP-built gaseous SMD prototype was tested with the AGS test beam in the energy range from ~1 to 7 GeV [7]. This SMD was positioned in the Small EMC prototype (SPEMC) at the depth of ~ $5X_o$. The beam Čerenkov counter was used to select electrons among other negative particles, mainly, π^- 's, striking SPEMC. In Fig.1 an electron shower shape in the SMD simulated with GEANT is compared with the experimental one. Their reasonable good agreement can be observed.



STAR barrel

Fig. 1. Experimental and Monte Carlo shower shapes in the SMD located after $5X_o$ within the EMC. Incident particles are 5 GeV electrons.

¹In the STAR EMC project, the anode wires were suggested to be paired to have almost equal, \sim 14-15 mm, segmentations in both directions. In this simulations, however, we followed the design of the SMD prototype of 1995 where each anode wire represented a separate readout channel.

The experimental shape was obtained [8] using two hodoscopes in front of SPEMC for measuring X- and Y-positions with 2-mm wide scintillator rods. Therefore, the experimental shower profile in Fig.1 is actually smeared with the 2-mm hodoscope resolution. The GEANT simulated shower shape for 5 GeV electrons was produced in a similar way with the same smearing which was actually achieved by 2-mm wide binning of the accumulated histogram for energy deposition in the SMD gas. For normalization, two profiles in Fig.1 have been equalized in their first bins.

Fig.2 shows the mean energy deposition in the SMD sensitive volume versus the energy of incident photons, for various SMD locations within the EMC. The energy deposited in the SMD is very small, of the order of $\approx 10^{-5}$ of the energy of incident photons or electrons. For the deep SMD position at $7X_o$, its response is quite linear up to $\sim 15-20$ GeV while at shallower locations, the noticeable non-linearity has shown up earlier.



STAR barrel

Fig. 2. Energy deposition in the SMD versus energy of incident photons for three SMD locations inside the EMC: $3X_o$, $5X_o$ and $7X_o$.

The GEANT simulated position resolutions in the SMD for γ -quanta are shown in Fig.3. In this figure, the characteristic σ is actually the "sigma" of the Gaussian fit to the distributions of differences between the "true" hit position and the "logarithmic weighted center of shower". For the "true" hit position, the crossing point of the γ 's trajectory with the SMD mid-plane was chosen. The coordinates of the "logarithmic weighted center of



Fig. 3. SMD position resolution for γ -quanta at various SMD locations in the EMC.

shower" in the SMD, X_c and Y_c , were calculated, using formulae [9]

$$X_{c} = \frac{\sum_{i} X_{i} W_{i}}{\sum_{i} W_{i}}; \quad Y_{c} = \frac{\sum_{i} Y_{i} W_{i}}{\sum_{i} W_{i}} \quad \text{with} \quad W_{i} = max\{0, W_{0} + ln(E_{i}/E_{tot})\}.$$
(1)

In the formulae above, E_i is the energy deposited under the i-th strip or wire; X_i is the position of the center line of the i-th strip and Y_i is the position of the i-th anode wire; E_{tot} is the total energy deposition in the SMD. The free parameter W_0 has been varied within some range to achieve the best results for the position resolution as well as for the recognition of multiple overlapping γ -hits which will be discussed later on in this paper. Among other things, the parameter W_0 effectively sets the threshold, floating from shower to shower, for the energy deposition in each particular channel, $E_{min} = E_{tot} \cdot \exp(-W_0)$.

A typical resolution in the SMD is in the scale of few millimetres. The resolution is better for the shallow SMD positions compared with deeper the locations. However, the spatial resolution is not the only parameter for the optimization of the SMD placement in the EMC. Other important characteristics such as γ and electron detection efficiencies, double-hit recognition, electron-hadron rejection, etc. must be also taken into account. The narrowing the SMD channel from 14 mm (strips) down to 7 mm (wires) gives almost proportional improvement in the spatial resolution. However, this is not "free" but requires the doubling of the number of the SMD channels as the defined in the project \sim 35–40k (barrel SMD only) to \sim 70–80k, and must also be weighted against the improvement of the EMC/SMD performance in general.

3. π^o/γ -Separation

In the simulations presented here, the effectiveness of the simple criteria for π^o/γ -separation based on the shower width in the SMD has been studied. These criteria were similar to those used in Refs.[10].

For each SMD shower, the "logarithmic weighted shower widths", $< R_x >$ and $< R_y >$, were calculated

$$< R_x > = \left[\frac{\sum_i W_i (X_i - X_c)^2}{\sum_i W_i}\right]^{1/2}; \quad < R_y > = \left[\frac{\sum_i W_i (Y_i - Y_c)^2}{\sum_i W_i}\right]^{1/2}.$$
 (2)

The meaning of parameters in Eqs.(2) is the same as in Eqs.(1) for the center of shower.

The single combined characteristic of the shower size, "mean weighted radius" $\langle R \rangle$, which we used for π^0/γ -separation was defined as

$$< R >= \sqrt{< R_x >^2 + < R_y >^2}.$$
 (3)

Typical $\langle R \rangle$ -distributions for the showers originated from single photons and π^{o} decays in the SMD at three their locations are shown in Figs. 4–6. At low energy end, two peaks, for single-photon and for π^{0} s, are clearly seen. In this energy range, the background from π^{0} s under the single- γ peak is mainly due to the events with one photon of two lost because of the detection inefficiency in the SMD. This inefficiency, along with other factors, significantly depends on the floating threshold set by the choice of parameter W_{0} (see Eqs.(1)). Therefore, setting the threshold at a feasibly lowest level is the key issue for catching as many as possible decays of low energy π^{0} s which is important to precisely evaluate the contribution of π^{0} s into the total γ -production in the collisions of interest and, eventually, to estimate the yield of direct photons with the acceptable accuracy.

At high energies, the distance in the SMD between two γ 's from π^{o} -decays becomes smaller, resulting in the fact that two showers usually overlap. As a consequence, the



Fig. 4. Typical distributions of the shower width, $\langle R \rangle$, for clusters, originating from single-photons and π^{o} -decays, in the SMD located at $3X_{o}$.

single-photon and π^0 separation at high energies is actually the problem of recognition and rejection of overlapping, presumably wider than single-hit clusters while keeping the efficiency to single photons, ϵ_{γ} , at a reasonably high level.

In this study, the $\langle R \rangle$ -cuts have been applied to separate single- γ and π^0 hits. These cuts along with the optimal values of parameter W_0 were chosen individually for each energy and SMD location from the requirement of having $\epsilon_{\gamma} = 80\%$.



Fig. 5. The same as in Fig.4 for the SMD located at $5X_o$.



Fig. 6. The same as in Fig.4 for the SMD located at $7X_o$.



Fig. 7. The simulated π^0 suppression power at 80% efficiency of single- γ detection for four SMD locations in the EMC stack.



Fig. 8. The comparison of π^0 suppression powers at 80% efficiency of single- γ detection for three SMD locations in the EMC stack.

The results of simulations of " π^0 suppression power"² are shown in Figs. 7 and 8. One can observe that the $\langle R \rangle$ -cut provides the highest suppression power for the SMD location at $\sim 5X_o$, and it works the best at the energies from ~ 3 to 10 GeV which overlap the range of interest for QGP search at RHIC through direct- γ yield measurements. Below and above this energy interval, the suppression power of $\langle R \rangle$ -cut drops quite significantly. This means that the study of others, more sophisticated criteria is necessary to improve γ/π^0 -separation at the both low and high ends of the energy spectrum of interest for the STAR experiment.

The "strengthening coefficient" of the signal-to-background ratio, that is $\epsilon_{\gamma}/(1-\epsilon_{\pi^o})$, is the largest at the energies ~3–7 GeV and, for the SMD location at $5X_o$, it reaches the value of $\simeq 4$ –6.5.

We have also tried to use the other characteristic of clusters in the SMD to recognize single- and double- γ hits, the "inertia moment" ER^2 which contains the information

²Which is actually the fraction of π^0 -hits truly recognized and rejected. i.e. the π^0 recognition efficiency, ϵ_{π^0} .

about the showers size weighted with the energy deposition

$$ER^2 = ER_x^2 + ER_y^2,\tag{4}$$

where $ER_x^2 = \sum_i E_i R_{xi}^2$ and $ER_y^2 = \sum_i E_i R_{yj}^2$ are the second moments for the strips and wires, respectively; $R_{xi} = X_c - X_i$ and $R_{yi} = Y_c - Y_i$. However, the simulations have shown that the criterion of "inertia moment" applied additionally to the $\langle R \rangle$ -cuts does not provide a noticeable improvement of γ/π^0 -separation in the SMD.

4. e/π -Rejection

The hadron suppression coefficient for the known particle momentum, K_R , (also, " e/π -rejection power") is defined here as a probability to misidentify a charged π -meson as an electron for 90% electron detection efficiency³. The $\sim 18X_o$ thick nonsegmented EMC provides the rejection at the level of K_R from 0.01 to 0.1 for the incident particle energies from 2 to 30 GeV. The < R >-cut for the shower width in the SMD gives an additional factor of ~ 2 -3, resulting in the total rejection power of the EMC+SMD of the order of 0.06–0.003. The results of simulations for e/π -rejection are presented in Fig.9. From these results, it also follows that the SMD position at $5X_o$ is more preferable than at $3X_o$.

Conclusion

Using the simplest method to distiguish single photons from $\pi^o \to 2\gamma$ decay, namely the shower width in the gaseous SMD, we have found that, in the wide energy range, the best results can be achieved for the SMD position close to $\sim 5X_o$ within the EMC stack. At this location for $E_{\gamma} \simeq 3-7$ GeV, the γ/π^o "signal-to-background" ratio can be improved by \sim 4–6.5 times. We hope that using some other EMC and SMD features, we can improve these numbers.

For example, at $P_T \leq 5-6$ GeV/c, two photons from π^0 -decay will mostly strike different EMC towers of size $\eta \times \varphi \simeq 0.05 \times 0.05$ and even SMD patches of the size 0.1×0.1 . In many cases, the energies and positions of both photons can be measured, and the invariant mass of the pair reconstructed. This provides the opportunity to evaluate the yield of low $P_T \pi^0$ -mesons and, eventually, the yield of direct photons in the high combinatorial background of the high multiplicity heavy-ion collisions. This is even more challenging task than the separation of two γ -hits which, at low P_T 's, are fairly far apart. The first simulation results, using this approach, [11,12] are quite encouraging, although

³To avoid a potential confusion, we have to underline that the coefficient K_R as it is defined here is rather the characteristic of the EMC+SMD themselves, which is usually measured in the beam tests [7], than the evaluation of the hadron suppression in the STAR detector. In the real setup, the particle momentum is never measured exactly. This will make the real hadron suppression power in STAR differ, sometimes significantly, from that provided here. This is particularly true for high P_T region where the momentum resolution in the STAR tracking system is considerably worse compared to the energy resolution in the EMC. The latter, in fact, was the only one of two taken here into account to set the cuts for achieving the 90% electron detection efficiency.



Fig. 9. e/π -rejection power vs energy of incident particles for two SMD locations in the EMC stack: at $3X_o$ and $5X_o$.

many details of the photon shower characteristics in the EMC and SMD have been omitted in these studies.

At the high energy end, more sophisticated algorithms for the recognition of single- γ and π^{o} -decays are to be explored (fit of shower shape, etc), including the possibility of using the power of neural networks (see, for example, Ref. [13]).

For the known charged particle momentum, the cuts on the energy deposition in the EMC around its mean value along with the cuts on the shower width in the SMD at $5X_o$ give the e/π -rejection power at the level of a few percent for the energies less than 5–10 GeV. For higher energies, the rejection power increases, and at 30 GeV, it reaches $\sim 3 \times 10^{-3}$. For the entire energy range from 2 to 30 GeV, the deep SMD positioning at $5X_o$ provides better e/π -rejection compared to the location at $3X_o$.

Acknowledgments

It's our pleasure to thank B. V. Chujko who kindly provided us with the measured shape of the electron shower in the SMD, and O. D. Tsai for the useful discussions.

References

- [1] S. A. Akimenko et al. Nucl. Instr. and Meth. A365 (1995) 92.
- [2] Proposal on Spin Physics Using the RHIC Polarized Collider (RHIC Spin Collaboration), August, 1992; Proposal Update, September, 1993.
- [3] C. P. Singh. Phys. Rev. **236** (1993) 147.
- [4] Xin-Nian Wang et al. Preprint LBL-38455.
- [5] P. K. Job et al. Nucl. Instr. and Meth. **A271** (1988) 442.
- [6] L. Balka et al. Nucl. Instr. and Meth. **A267** (1988) 272.
- [7] W. J. Llope (for the STAR-EMC Collaboration). Report at the "VI International Conference on Calorimetry in High Energy Physics", INFN, Frascati (Roma), Italy, June 8–14, 1996; Proc. in Frascati Physics Series, Volume VI, Eds. A. Antonelli, S. Bianco, A. Calcaterra, F. L. Fabbri. pp.187–197.
- [8] B. V. Chujko. Private Communication.
- [9] T. C. Awes et al. Nucl. Instr. and Meth. A311 (1992) 130.
- [10] J. Grunhaus et al. Preprint TAUP-1976-92, Tel Aviv University, Tel-Aviv, Israel, 1992.
- [11] T. M. Cormier. Report at the STAR Collaboration Meeting, August 1996, BNL (unpublished).
- [12] T. J. LeCompte. STAR Note #251, July 1996.
- [13] N. G. Minaev. Preprint IHEP 94–142, Protvino, 1994.

Received July 10, 1997

А.А.Деревщиков, Ю.А.Матуленко, В.Л.Рыков и др.

Расчеты эффективности γ/π^{o} -разделения и e/h-реакции в центральном электромагнитном калориметре и детекторе максимума ливня установки STAR.

Оригинал-макет подготовлен с помощью системы ЦАТ _Е Х.	
Редактор Е.Н.Горина.	Технический редактор Н.В.Орлова.
Подписано к печати 18.07.97	. Формат $60 \times 84/8$. Офсетная печать.
Печ.л. 1.75. Учизд.л. 1.34.	Тираж 180. Заказ 1140. Индекс 3649.
ЛР №020498 17.04.97.	

ГНЦ РФ Институт физики высоких энергий 142284, Протвино Московской обл.

Индекс 3649

 $\Pi P Е \Pi P И H Т 97-49,$ $И \Phi В Э,$ 1997